# **Measuring particle energy**

# Giovanni De Lellis

University Federico II and INFN Naples, Italy NUST MISÍS, Moscow, Russia

Giovanni.de.lellis@cern.ch



# **Calorimetry in Particle Physics**

This lecture draws heavily from the Review Article 'Calorimetry for Particle Physics', C.W. Fabjan and F. Gianotti, Rev. Mod. Phys., Vol. 75, NO. 4, October 2003

Much information was also taken from the massive Monograph

'Calorímetry, Energy Measurement in Particle Physics', R. Wigmans, Oxford University Press, Second edition, 2017

Training lectures at CERN by Werner Riegler



#### Bremsstrahlung

A charged particle of mass M and charge  $q = Z_1 e$  is deflected by a nucleus of charge Ze which is partially 'shielded' by the electrons. During this deflection the charge is 'accelerated' and it therefore radiates  $\rightarrow$  Bremsstrahlung.

From Bethe's theory the elastic scattering off the Nucleus is given by

$$\epsilon_0(q) = Z_2 - \sum_{j=1}^{Z_2} \int e^{i(\vec{q}\vec{r}_j)} \psi_0^2(\vec{r}_j) d^3 r_1 \dots d^3 r_{Z_2} = Z_2 - F \qquad \frac{d\sigma}{d\Omega} = \left(\frac{1}{4\pi\varepsilon_0} \frac{Z_1(Z_2 - F)e_0^2}{2pv}\right)^2 \frac{1}{\sin^4\theta/2}$$

where F(q) describes the partial shielding of the nucleus by the electrons. Effective values for F are used in the following expressions.



$$\frac{d\sigma}{d\Omega} = \left(\frac{2Z_1Z_2e^2}{4\pi\epsilon_0 \ p \cdot v}\right)^2 \frac{1}{(2\sin\frac{\theta}{2})^4} \qquad p = Mv\gamma$$
"Rutherford Scattering"
Written in Terms of Momentum Transfer:  $Q^2 = 2p^2(1 - q)$ 

$$\frac{d\sigma}{dQ} = 8\pi \left(\frac{Z_1Z_2e^2}{4\pi\epsilon_0\beta c}\right)^2 \cdot \frac{1}{Q^2} \qquad Q = |\vec{p} - \vec{p}'|$$

$$\lim_{\omega \to 0} \frac{dI}{d\omega} \sim \frac{2}{3\pi} \frac{Z_1^2e^2}{M^2c^3} \frac{Q^2}{4\pi\epsilon_0} \text{ Radiated energy between } \boldsymbol{\omega} \text{ and } \boldsymbol{\omega}'$$

$$\frac{dE}{dx} = \frac{N_A\rho}{A} \cdot \int_0^{\omega_{max}} d\omega \int_{Q_{min}}^{Q_{max}} dQ \frac{dI}{d\omega} \cdot \frac{d\sigma}{dQ} \qquad \omega_{max} = \frac{E}{\hbar}$$

$$\frac{dE}{dx} = \frac{N_A\rho}{A} \cdot \frac{16}{3}\alpha \cdot Z^2 \cdot \left(\frac{Z_1^2e^2}{4\pi\epsilon_0Mc^2}\right)^2 \cdot E \cdot \ln\frac{Q_{max}}{Q_{min}}$$

$$\alpha = \frac{e^2}{4\pi\epsilon_0\hbarc} \sim \frac{1}{137}$$

OPENING A NEW CENTURY Science and Technology

A charged particle of mass M and charge  $q = Z_1 e$  is deflected by a nucleus of Charge Ze.

Because of the acceleration  $-\cos\theta$ ) the particle radiates EM waves - energy loss.

> coulomb-scattering (Rutherford Scattering) describes the deflection of the particle.

Maxwell's Equations describe the radiated energy for a given momentum transfer.

 $\rightarrow dE/dx$ 

# Bremsstrahlung: Quantum Mechanics

q,M,E-E'

7

$$q = Z_1 e_1, \ E + Mc^2 >> 137Mc^2 Z^{-\frac{1}{3}}$$

→ Highly Relativistic:

$$\frac{d\sigma(E,E')}{dE'} = 4\alpha Z^2 Z_1^4 \left(\frac{1}{4\pi\epsilon_0}\frac{e^2}{Mc^2}\right)^2 \frac{1}{E'}F(E,E')$$

$$F(E,E') = \left[1 + \left(1 - \frac{E'}{E+Mc^2}\right)^2 - \frac{2}{3}\left(1 - \frac{E'}{E+Mc^2}\right)\right] \ln 183Z^{-\frac{1}{3}} + \frac{1}{9}\left(1 - \frac{E'}{E+Mc^2}\right)$$

$$\frac{dE}{dx} = -\frac{N_A\rho}{A} \int_0^E E' \frac{d\sigma}{dE'} dE' \sim 4\alpha Z^2 Z_1^4 \left(\frac{1}{4\pi\epsilon_0}\frac{e^2}{Mc^2}\right)^2 E \left[\ln 183Z^{-\frac{1}{3}} + \frac{1}{1}\frac{1}{4\pi\epsilon_0}\frac{e^2}{Mc^2}\right]$$

$$dx = A^{-1} (4\pi\epsilon_0 Mc^2)^{-1} E^{-1} (1002)$$

$$E(x) = E_0 \exp^{-\frac{x}{X_0}} A^{-1}$$

X<sub>0</sub> radiation length





 $4\alpha N_A \rho Z^2 \left(\frac{1}{4\pi\epsilon_0} \frac{e^2}{Mc^2}\right)^2 \ln\left(183Z^{-\frac{1}{3}}\right)$ 

Proportional to  $Z^2/A$  of the Material.

Proportional to  $Z_1^4$  of the incoming particle.

Proportional to p of the material.

Proportional 1/M<sup>2</sup> of the incoming particle.

Proportional to the Energy of the  $\frac{1}{8}$  | Incoming particle  $\rightarrow$ 

 $E(x)=e(-x/X_0) - 'Radiation Length'$ 

 $X_0 \propto M^2 A / (\varrho Z_1^4 Z^2)$ 

 $X_0$ : Dístance where the Energy  $E_0$ of the incoming particle decreases  $E_0 e^{-1} = 0.37 E_0$ 

### **Critical Energy**

such as copper to about 1% accuracy for energies between about 6 MeV and 6 GeV



For the muon, the second lightest particle after the electron, the critical energy is at 400GeV.

The EM Bremsstrahlung is much less important for muons: at the LHC and in cosmic-rays experiments can be relevant

Crítical Energy: If dE/dx (Ionízation) = dE/dx (Bremsstrahlung)

Muon in Copper:  $p \approx 400 \text{GeV}$ 

Electron in Copper:  $p \approx 20 \text{MeV}$ 



NUST MISIS, Russia, Moscow

#### **Pair Production: Quantum Mechanics**



The diagram is very similar to Bremsstrahlung

Crossing Symmetry: same cross-section For  $E\gamma \!\!>\!\! m_ec^2 \!\!=\!\! 0.5 MeV: \lambda = 9/7X_0$ 

Average distance a high energy photon has to travel before it converts into an e<sup>+</sup>e<sup>-</sup> pair is equal to 9/7 of the distance that a high energy electron has to travel before reducing its energy from  $E_0$  to  $E_0^*e^{-1}$  by photon radiation.



#### **Bremsstrahlung + Pair Production — EM Shower**



Electromagnetic Shower → EM Calorimeter



¥ S]

#### Electro-Magnetic Shower of High Energy Electrons and Photons

Number of particles  $(e^{\pm}, \gamma)$  after  $nX_0$ 

Average Energy of particles after  $nX_{0}$ 

Shower stops if  $E(n) = E_c$ 

 $n_{max} = rac{1}{\ln 2} \ln rac{E_0}{E_c}$  shower length rises with  $\ln E_0$ 

Number of  $e^{\pm}$  track segments of length  $X_0$  after  $nX_0$ 

$$L = \sum_{n=0}^{n_{max}} 2^n X_0 = \left(2\frac{E_0}{E_c} - 1\right) X_0 \sim 2\frac{E_0}{E_c} X_0 = c_1 \cdot E_0$$

Total (charged) track length is proportional to the Energy of the particle → Calorimeter Principle



 $N(n) = 2^n$ 

 $E(n) = \frac{E_0}{2^n}$ 

 $N_{tr}(n) = 2^n$ 

NUST MISIS, Russia, Moscow

# Calorimetry: Energy Measurement by total Absorption of Particles



The  $e^{\pm}$  in the Calorimeter ionize and excite the material

Ionization: e-, I+ pairs in the material

Excitation: Photons in the material

Measuring the total number of e-, I+ pairs or the total number of photons gives the particle Energy







If N is the total number of  $e^{\text{-}},\,I^{+}$  pairs or photons,  $N=c_{1}E_{0};$ 

 $\Delta N=\sqrt{N}$  (Poisson statistics)

 $\frac{\Delta E}{E} = \frac{\Delta N}{N} = \frac{1}{\sqrt{N}} = \frac{a}{\sqrt{E}}$  Radiation

Only Electrons and High Energy Photons show EM cascades at current GeV-TeV level Energies.

Strongly interacting particles like Pions, Kaons, produce hadronic showers in a similar fashion to the EM cascade → Hadronic calorimetry

# Calorimetry: Energy Measurement by total Absorption of Particles

The measurement is destructive. The particle cannot be studied further



11

# Calorimetry

Calorímeters are blocks of instrumented material in which particles to be measured are fully absorbed and their energy transformed into a measurable quantity.

The interaction of the incident particle with the detector (through electromagnetic or strong processes) produces a shower of secondary particles with progressively degraded energy.

The energy deposited by the charged particles of the shower in the active part of the calorimeter, which can be detected in the form of charge or light, serves as a measurement of the energy of the incident particle.



C.W. Fabjan and F. Gianotti, Rev. Mod. Phys., Vol. 75, N0. 4, October 2003

# Calorimetry

Calorímeters can be classified into:

- Electromagnetic Calorimeters, to measure electrons and photons through their EM interactions.
- Hadron Calorímeters,

used to measure hadrons through their strong and EM interactions.

The construction can be classified into:

- Homogeneous Calorímeters, that are built of only one type of material that performs both tasks, energy degradation and signal generation.
- Sampling Calorimeters,

that consist of alternating layers of an absorber, a dense material used to degrade the energy of the incident particle, and an active medium that provides the detectable signal.







C.W. Fabjan and F. Gianotti, Rev. Mod. Phys., Vol. 75, N0. 4, October 2003

# Calorimetry

Calorímeters are attractive in our field for various reasons:

In contrast with magnet spectrometers, in which the momentum resolution deteriorates linearly with the particle momentum, on most cases the calorimeter energy resolution improves as 1/Sqrt(E), where E is the energy of the incident particle. Therefore calorimeters are very well suited for high-energy physics experiments.

In contrast to magnet spectrometers, calorímeters are sensitive to all types of particles, charged and neutral. They can even provide indirect detection of neutrinos and their energy through a measurement of the event missing energy.

Calorímeters are commonly used for trígger purposes sínce they can províde fast sígnals that are easy to process and interpret.

They are space and therefore cost effective. Because the shower length increases only logarithmically with energy, the detector thickness needs to increase only logarithmically with the energy of the particles. On the contrary, for a fixed momentum resolution, the bending power  $BL^2$  of a magnetic spectrometer must increase linearly with the particle momentum.







C.W. Fabjan and F. Gianotti, Rev. Mod. Phys., Vol. 75, N0. 4, October 2003

#### **Interaction of Particles with Matter**

Any device to detect a particle must interact with it in some way -> almost ...

- Neutrínos can be measured by missing transverse energy.
- E.g. p p collider  $E_T=0$ ,
- If the  $\Sigma E_T$  of all collision products is  $\neq 0 \rightarrow$  neutrino escaped



# **EM Calorimetry**

Approximate longitudinal shower development

 $N(n) = 2^n$  Number of particles (e<sup>±</sup>,  $\gamma$ ) after nX<sub>0</sub>

 $E(n) = rac{E_0}{2^n}$  Average Energy of particles after  $nX_0$ 

Shower stops if  $E(n)=E_{\rm c}$ 

 $n_{max} = rac{1}{\ln 2} \ln rac{E_0}{E_c}$  shower length rises with  $\ln E_0$ 

Radiation Length  $X_0$  and Moliere Radius are two key parameters for the choice of calorimeter materials



#### Approximate transverse shower development

The traverse shower direction is mainly related to the Multiple Coulomb Scattering of the low Energy Electrons

$$\begin{split} \theta_0 &\sim \frac{21(mrad)}{\beta p \ (MeV)} Z_1 \sqrt{\frac{X}{X_0}} \\ \textbf{Electrons:} \quad E_c, \ E &\sim p \cdot c \\ \theta_0 &\sim \frac{21(mrad)}{\beta E_c \ (MeV)} Z_1 \sqrt{\frac{X}{X_0}} \qquad Z_1 = 1, \ \beta \sim 1 \\ E_c &\sim \frac{610}{Z+1.24} MeV \sim \frac{610}{Z} MeV \\ \theta &= 0.0344 \cdot Z \cdot \sqrt{\frac{X}{X_0}} \end{split}$$

Molíere Radíus  $\varrho_m$  = Lateral Shower Radíus after  $1X_0$ 

 $\rho_m \approx 0.0344 \cdot Z \cdot X_0$ 

95% of Energy is in a Cylinder of  $2\varrho_m$  radius

# Simulated EM Shower Profiles in PbWO<sub>4</sub>



FIG. 2. (a) Simulated shower longitudinal profiles in  $PbWO_4$ , as a function of the material thickness (expressed in radiation lengths), for incident electrons of energy (from left to right) 1 GeV, 10 GeV, 100 GeV, 1 TeV. (b) Simulated radial shower profiles in  $PbWO_4$ , as a function of the radial distance from the shower axis (expressed in radiation lengths), for 1 GeV (closed circles) and 1 TeV (open circles) incident electrons. From Maire (2001).

In calorimeters with thickness  $\sim 25 \text{ X}_0$ , the shower leakage beyond the end of the active detector is much less than 1% up to incident electron energies of  $\sim 300 \text{ GeV}$  (LHC energies).



## **Crystals for Homogeneous EM Calorimetry**

In crystals the light emission is related to the crystal structure of the material. Incident charged particles create electron-hole pairs and photons are emitted when electrons return to the valence band.

The incident electron or photon is completely absorbed and the produced amount of light, which is reflected through the transparent crystal, is measured by photomultipliers (PM) or solid state photon detectors (SiPM)



Measuring the Photons produced by the collision of the e<sup>±</sup> with atom electrons of the material



### Crystals for Homogeneous EM Calorimetry

	NaI(Tl)	CsI(Tl)	CsI	BGO	PbWO <sub>4</sub>
Density (g/cm <sup>3</sup> )	3.67	4.53	4.53	7.13	8.28
$X_0$ (cm)	2.59	1.85	1.85	1.12	0.89
$R_M$ (cm)	4.5	3.8	3.8	2.4	2.2
Decay time (ns)	250	1000	10	300	5
slow component			36		15
Emission peak (nm) slow component	410	565	305 480	410	440
Light yield $\gamma$ /MeV	$4 \times 10^{4}$	$5 \times 10^{4}$	$4 \times 10^{4}$	$8 \times 10^{3}$	$1.5 \times 10^{2}$
Photoelectron yield (relative to NaI)	1	0.4	0.1	0.15	0.01
Rad. hardness (Gy)	1	10	$10^{3}$	1	$10^{5}$

KTeV@Tevat L3@LEP, CMS@LHC, Babar@PEPII, 25ns bunch 10ms interaction<sup>ron,</sup> 25**u**s High rate, bunch crossing, rate, good light Good crossing, high yield, good S/N resolution radiation Low radiation dose

dose





SHiP



#### **Crystals for Homogeneous EM Calorimetry**





Fig. 2. Longitudinal drawing of module 2, showing the structure and the front-end electronics layout.



#### Noble Liquids for Homogeneous EM Calorimetry

	Ar	Kr	Xe
Z	18	36	58
Α	40	84	131
$X_0$ (cm)	14	4.7	2.8
$R_M$ (cm)	7.2	4.7	4.2
Density (g/cm <sup>3</sup> )	1.4	2.5	3.0
Ionization energy (eV/pair)	23.3	20.5	15.6
Critical energy $\epsilon$ (MeV)	41.7	21.5	14.5
Drift velocity at saturation $(mm/\mu s)$	10	5	3



when a charge particle traverses these materials, about half the lost energy is converted into ionization and half into scintillation.

The best energy resolution would obviously be obtained by collecting both the charge and light signal. This is however rarely done because of the technical difficulties to extract light and charge in the same instrument.

Krypton is preferred in homogeneous detectors due to small radiation length and therefore compact detectors. Liquid Argon is frequently used due to low cost and high purity in sampling calorimeters.







#### NUST MISIS, Russia, Moscow Noble Liquids for Homogeneous EM Calorimetry



SHi

Science and Technolog

22

#### Homogeneous EM Calorimeters, Examples

NA48 Liquid Krypton 2cmx2cm cells  $X_0 = 4.7$ cm 125cm length (27 $X_0$ )  $\rho = 5.5$ cm





Fig. 1. Schematic of the KTeV CsI Calorimeter showing the cluster energy profiles due to four photon

NA48 Experiment at CERN and KTeV Experiment at Fermilab, both built for measurement of direct CP violation. Homogenous calorimeters with Liquid Krypton (NA48) and CsI (KTeV). Excellent and very similar resolution.







# **Energy Resolution of Calorimeters**

Stochastic term: Fluctuations related to the physics development of the shower.

σ

 $\boldsymbol{E}$ 

Noise term:

From electronics noise of the readout chain. For constant electronics noise  $\rightarrow$  double signal = double S/N

Constant term: Instrumental effects that cause variations of the calorimeter response with the particle impact point.

Add in squares

For homogeneous calorímeters the noise term and constant term become dominant.

For sampling calorimeters the stochastic term, then called 'sampling' term becomes dominant.



FIG. 3. Fractional electron energy resolution as a function of energy measured with a prototype of the NA48 liquid krypton electromagnetic calorimeter (NA48 Collaboration, 1995). The line is a fit to the experimental points with the form and the parameters indicated in the figure.



# **Sampling Calorimeters**



Alternation of "passive" absorber plates and "active" readout sections

Advantage:

- optímum choíce of absorber materíal
- optimum choice of signal readout
- compact and cheap construction
   "passive": Pb, Fe...

"active": Scintillator (signal → Photons) Noble Liquid, e.g. Ar (Signal → e-, I-) Wire Chambers (Signal → e-, I+) Emulsions (Signal --> tracks)

Energy resolution of sampling calorimeters is in general worse than that of homogeneous calorimeters, owing to the sampling fluctuations – the fluctuation of ratio of energy deposited in the active and passive material.

The resolution is in the range 5-20%/Sqrt[E(GeV)] for EM calorimeters. On the other hand they are relatively easy to segment longitudinally and laterally and therefore they usually offer better space resolution and particle identification than homogeneous calorimeters.

Active medium can be scintillators, solid state detectors, emulsion films, gas detectors or liquids. Sampling Fraction = Energy deposited in Active/Energy deposited in passive material.







# Gas and Solid State Sampling Calorimeters



Gas sampling calorimeters have been widely employed at LEP because of their low cost and segmentation flexibility.

They are not well suited to present and future machines because of their modest EM energy resolution ~ 20%/Sqrt[E(GeV)].

Solid state detectors as active readout medium use mostly silicon. The advantage is very high signal to noise ratio (large signals). Often used on a small scale as luminosity monitors.

The disadvantage is the high cost, preventing large calorimeters, and modest radiation resistance.



# **Scintillator Sampling Calorimeters**



Wavelength shifters absorb photons from the scintillators and emit light at a longer wavelength which does not go back into the scintillator but is internally reflected along the readout plate to the photon detector ->

the readout plate to the photon detector → compact design.

A large number of sampling calorimeters use organic scintillators arranged in fibers or plates.

The drawbacks are that the optical readout suffers from radiation damage and nonuniformities at various stages are often the source of a large constant term.



KLOE EM calorímeter:  $\frac{5\%}{\sqrt{E(GeV)}}$ 

FIG. 13. Schematic layout of the barrel part of the KLOE electromagnetic calorimeter (Antonelli et al., 1995).



# **Liquid Sampling Calorimeters**



FIG. 15. Schematic view of a traditional sampling calorimeter geometry (a) and of the accordion calorimeter geometry (b).



MISIS National University of Science and Technology



Líquíds at room temperature do not requíre cryogeny but are characterízed by poor radiation resistance and suffer from purity problems

 $\rightarrow$ Noble líquíds at cryogeníc temperatures.

The advantages are operation in 'ion chamber mode', i.e. deposited charge is large and doesn't need multiplication, which ensures better uniformity compared to gas calorimeters that need amplification.

They are relatively uniform and easy to calibrate because the active medium is homogeneously distributed inside the volume. They provide good energy resolution (e.g. ATLAS 10%/Sqrt[E(GeV)])

And stable operation with time.

They are radiation hard.

with the standard liquid argon sampling calorimeters the alternating absorber and active layers are placed perpendicular to the direction of the incident particle.

→ Long cables are needed to gang together the readout electrodes, causing signal degradation, dead spaces between the calorimeter towers and therefore reduced hermeticity.

# **Liquid Argon Sampling Calorimeters**



FIG. 15. Schematic view of a traditional sampling calorimeter geometry (a) and of the accordion calorimeter geometry (b).

For the ATLAS LAr Calorímeter this was solved by placing the absorbers in an accordion geometry parallel to the particle direction and the electrodes can be read out from the 'back side'.

ATLAS: Lead layers of 1.1-2.2mm, depending on the rapidity region, separated by 4mm liquid Argon gaps.

Test beam results have shown  $10\%/Sqrt[E(GeV)] \times 0.25/E(GeV) \times 0.3\%$ 





FIG. 17. Schematic view of the segmentation of the ATLAS electromagnetic calorimeter.

NUST MISIS, Russia, Moscow

# Emulsion Cloud Chamber: a peculiar type of sampling calorimeter



FROM MGA TO NUST MIS

# Electromagnetic showers as seen in an Emulsion Cloud Chamber: $e/\pi^0$ separation



High sampling frequency, 6 active layers every X<sup>0</sup>

 $\rightarrow$  high spatial resolution  $\rightarrow$  high purity in  $e/\pi^0$  separation



### Electron neutrino interaction with a $\pi^0$

#### Transverse plane









# Algorithms for shower and energy reconstruction



Number of the tracks inside the cone for a given opening angle

Total Number of tracks related to the shower



Efficency =

# Variables Used to discard the Instrumental Background



FROM MGA TO NUST MI

#### **Different performance in different zones**



#### Variables used in the analysis of Zone 1 events

21 plate A cone with opening angle 50 mrad is 3 cm defined, starting from the decay point for the  $\tau \rightarrow e$  decays, and from the 3 cm primary vertex for ve events.  $\alpha_{max} = 50 \text{ mrad}$ 57 Plate Input variable: IPA Input variable: alfa Input variable: deltaSx Signal (1/N) dN/0.00128 1/N) dN/0.02 Background /N) dN/0.04 50 4 40 30 20 0.01 0.02 0.03 0.04 0.05 0.2 0.4 0.6 0.8 -0.8 -0.6 -0.4 -0.2 0 0.2 0.4 0.6 0.8 IPA alfa deltaSx Input variable: deltaSy Input variable: chi2 3240.0/Nb (N/1) (1/N) dN/0.062 0.8 0.6 0.4 0.2 -0.8 -0.6 -0.4 -0.2 0 0.2 0.4 0.6 0.8 0.5 1.5 2.5 1 2 deltaSy chi2 FROM MGA TO NUST MIS



National University of

Science and Technology

2018

SHil

### **Energy Resolution**



Average number of BTs per event (Signal) = 160Average number of BTs per event (BG) = 17



#### Analysis extended to the downstream brick



#### **Resolution using the downstream brick**



# **Hadronic Calorimetry**

Hadron  $p^{\pm}, n, \pi^{\pm}, K^{\pm}, K^0$ 



Strong Interaction

Approximate Energy Distribution

$\sim$	50%	)	
$\pi^+, \tau$	τ_	$\pi^0$	
	$\backslash$	Ļ	
	. +	<u>ΥΥ</u>	0
$\pi^+,\pi^-$	$\pi$	$,\pi$	$,\pi$
\	/ \	/ \	$\gamma \gamma$
Hadro	n C	asci	ade

~20% Nuclear Excitation 5-30 MeV p, n, γ

~30% Slow Nucleons

 $\pi^0 \rightarrow \gamma \gamma \rightarrow \text{Electromagnetic Component}$ 



In Hadronic Cascades the longitudinal Shower is given by the Absorption Length  $\lambda_{\rm a}$   $I\sim \exp^{-rac{x}{\lambda_a}}$ 

In typical Detector materials  $\lambda_a$  is much larger than  $X_0$   $\lambda \sim \frac{1}{9} \cdot 35A^{\frac{1}{3}}$   $\rho$  X<sub>0</sub>  $\lambda$ Fe 7,87 1.76cm ~17cm

**Pb** 11,35 0.56cm ~17cm

Energy Resolution:

- A large fraction of the Energy "disappears" into:
  - Binding Energy of the emitted Nucleons
  - $\pi \rightarrow \mu + \nu$  which are not absorbed
- $\pi^{0}$  's decaying into  $\gamma\gamma$  start EM Cascade  $(\tau{\sim}10^{\text{-16}}\text{s})$

Energy resolution is worse than for EM Calorimeters

NUST MISIS, Russia, Moscow

# Hadron Calorimeters are Large because $\lambda$ is large



41

Hadron Calorímeters are large and heavy because the hadronic interaction length  $\lambda$ , the "strong interaction equivalent" to the EM radiation length  $X_0$ , is large (5-10 times larger than  $X_0$ )



![](_page_40_Picture_5.jpeg)

By analogy with EM showers, the energy degradation of hadrons proceeds through an increasing number of (mostly) strong interactions with the calorimeter material.

However the complexity of the hadronic and nuclear processes produces a multitude of effects that determine the functioning and the performance of practical instruments, and make hadronic calorimeters more complicated instruments to optimize.

The hadronic interaction produces two classes of effects:

Fírst, energetic secondary hadrons are produced. Their momenta are typically a sizable fraction of the primary hadron momentum i.e. at the GeV scale.

Second, in hadronic collisions with the material nuclei, a significant part of the primary energy is consumed in nuclear processes such as excitation, nucleon evaporation, spallation etc., resulting in particles with characteristic nuclear energies on the MeV scale.

Because part of the energy is therefore 'invisible', the resolution of hadron calorimeters is typically worse than in EM calorimeters 20-100%/Sqrt[E(GeV)].

![](_page_41_Picture_8.jpeg)

![](_page_42_Figure_2.jpeg)

# 'Deciphering this message becomes the story of hadronic calorimetry'

C.W. Fabjan and F. Gianotti, Rev. Mod. Phys., Vol. 75, No. 4, October 2003

FIG. 19. Particle spectra produced in the hadronic cascade initiated by 100-GeV protons absorbed in lead. The energetic component is dominated by pions, whereas the soft spectrum is composed of photons and neutrons. The ordinate is in "lethargic" units and represents the particle track length, differential in log *E*. The integral of each curve gives the relative fluence of the particle. Fluka calculations (Ferrari, 2001).

The signals from an electron or photon entering a hadronic calorimeter is typically larger than the signal from a hadron cascade because the hadronic interactions produce a fair fraction of invisible effects (excitations, neutrons ...)

![](_page_43_Figure_3.jpeg)

Because a fair fraction of shower particles consists of  $\pi^0$  which instantly decay into two photons, part of the hadronic cascade becomes an EM cascade – 'and never comes back'.

Because the EM cascade had a larger response than the Hadron cascade, the event/event fluctuation of produced  $\pi^0$  particles causes a strong degradation of the resolution.

Is it possible to build a calorimeter that has the same response (signal) for a 10GeV electron and 10GeV hadron ?  $\rightarrow$  compensating calorimeters.

![](_page_44_Figure_5.jpeg)

### **Compensating Hadron Calorimeters**

In a homogeneous calorímeter ít ís clearly not possíble to have the same response for electrons and hadrons.

For sampling calorimeters the sampling frequency and thickness of active and passive layers can be tuned such that the signal for electrons and hadrons is indeed equal ! Using Uranium or Lead with scintillators, hadron calorimeters with excellent energy resolution and linearity have been built.

Energy resolution

![](_page_45_Figure_5.jpeg)

### **Compensating Hadron Calorimeters**

Resolution and linearity of a hadron calorimeter is best if e/h=1. For all other values, the resolution in linearity is worse.

![](_page_46_Figure_3.jpeg)

![](_page_46_Picture_4.jpeg)

![](_page_46_Figure_5.jpeg)

FIG. 24. Monte Carlo simulation of the effects of  $e/\pi \neq 1$  on energy resolution (a) and response linearity (b) of hadron calorimeters with various values for e/h (intrinsic), where h(intrinsic) denotes the response to the purely hadronic component of the shower (Wigmans, 1988).

![](_page_46_Picture_7.jpeg)

### **Particle ID**

![](_page_47_Figure_2.jpeg)

C. Lippmann - 2003

![](_page_47_Picture_4.jpeg)

#### **Cherenkov Radiation**

If the velocity of a charged particle is larger than the velocity of light in the medium v>c/n (n = Refractive index of material) it emits Cherenkov radiation of a characteristic angle of  $\cos\theta_c = 1/n\beta$  ( $\beta = v/c$ )

$$\frac{dN}{dx} \sim 2\pi\alpha Z_1^2 \left(1 - \frac{1}{\beta^2 n^2}\right) \frac{\lambda_2 - \lambda_1}{\lambda_2 \cdot \lambda_1}$$

Numbers of emitted Photons\length with  $\lambda$  between  $\lambda_1$  and  $\lambda_2$  with  $\lambda_1$ =400 nm,  $\lambda_2$ =700 nm

$$\frac{dN}{dx} = 490 \left(1 - \frac{1}{\beta^2 n^2}\right) \left[\frac{1}{cm}\right]$$

![](_page_48_Picture_6.jpeg)

Material	n-1	β threshold	γ threshold
solid sodium	3,22	0,24	1,029
lead glass	0,67	0,60	1,25
water	0,33	0,75	1,52
silica aerogel	0.025-0.07 5	0.93-0.976	2.7-4.6
air	2.99.10-4	0,9997	41,2
He	3.3.10-5	0,99997	123

#### **Ring Imaging Cherenkov Detector**

![](_page_49_Figure_2.jpeg)

1000 200 GeV/c K,  $\pi$ 100 10 1 65 66 67 68 69 70 Ring Radius [mm]

Resolution:

![](_page_49_Figure_5.jpeg)

![](_page_49_Picture_6.jpeg)

![](_page_49_Picture_7.jpeg)

SHil

![](_page_49_Picture_8.jpeg)

![](_page_49_Figure_9.jpeg)

![](_page_49_Figure_10.jpeg)

![](_page_49_Figure_11.jpeg)

#### Measuring the neutrino energy: the OPERA experiment as an example

![](_page_50_Figure_2.jpeg)

![](_page_50_Picture_3.jpeg)

### **Measuring neutrino energy in OPERA**

![](_page_51_Figure_2.jpeg)

![](_page_51_Figure_3.jpeg)

![](_page_51_Figure_4.jpeg)

#### $\nu_{\mu}CC$ events with at least one reconstructed muon

![](_page_51_Figure_6.jpeg)

FROM MGA TO NUST MISIS

1918-2018

ENING A NEW CENTURY

### **Measuring neutrino energy in OPERA**

![](_page_52_Figure_2.jpeg)

NUST MISIS, Russia, Moscow

#### **Measuring neutrino energy in OPERA**

![](_page_53_Figure_2.jpeg)

MINGA TO NUST MISS 000 g 018-2018 DISA REV CENTURY DISA NEW CE

Total reconstructed energy for events with at least one identified muon for data (dots with error bars) and MC (solid line). The MC distribution is normalized to data.

$$y_B = 1 - \frac{E_\mu}{E_{\nu_\mu}} = \frac{E_{had}}{E_\mu + E_{had}}$$

Bjorken-y variable reconstructed in data (dots with error bars) and MC (shaded areas) for all the events with at least one muon. The MC distributions are normalized to data.

# Measuring energy with a SHiP prototype

 $\sigma(E_{had}) = (18.8 \pm 0.2)\%$ 

- Prototype of the SHíP Scattering and Neutrino detector to take data at the LHC in Runz
- Combining electromagnetic (Emulsion + SciFi) and hadronic calorimeter (SciFi +

![](_page_54_Picture_4.jpeg)

![](_page_54_Picture_5.jpeg)

![](_page_54_Figure_6.jpeg)

Resolution

![](_page_54_Figure_8.jpeg)

### Conclusion

The story of modern calorimetry is a textbook example of physics research driving the development of an experimental method.

The long quest for precision electron and photon spectroscopy explains the remarkable progress in new instrumentation techniques, for both sampling and homogeneous calorimeters.

The study of jets of particles as the macroscopic manifestation of quarks has driven the work on hadronic calorimeters.

Calorímeters largely used also in neutrino physics

![](_page_55_Picture_6.jpeg)